



ARTEMIS: A New Mission Using THEMIS B & C Spacecraft

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Agenda

- THEMIS Mission Overview
- THEMIS/ARTEMIS Mission Extension
- ARTEMIS Concept of Operations
- ARTEMIS Science
- ARTEMIS Schedule







Time History of Events and Macroscale Interactions during Substorms (THEMIS)

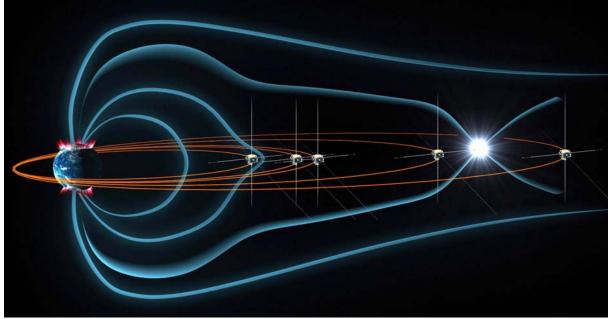
NASA Medium-class explorer mission, managed in PI Mode by UCB/UCLA

PI: Dr. Vassilis Angelopoulos

http://themis.ssl.berkeley.edu

5 Probes in synchronized Earth orbits with periods of 4, 2 and 1 days

Conjunctions are formed in magnetospheric tail every 4 days

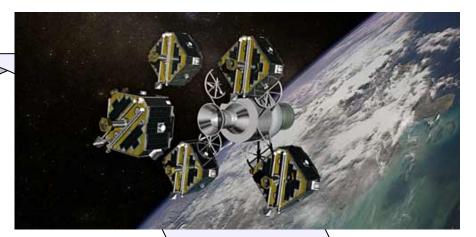








Launch: February 17, 2007



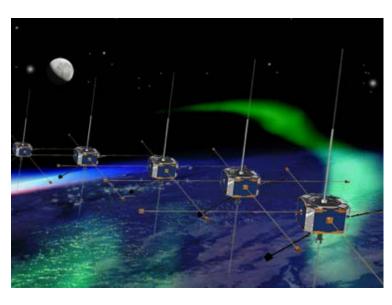
Probe release: L + 70 min



Responsibility for constellation operations transitioned to **MOC** at UCB

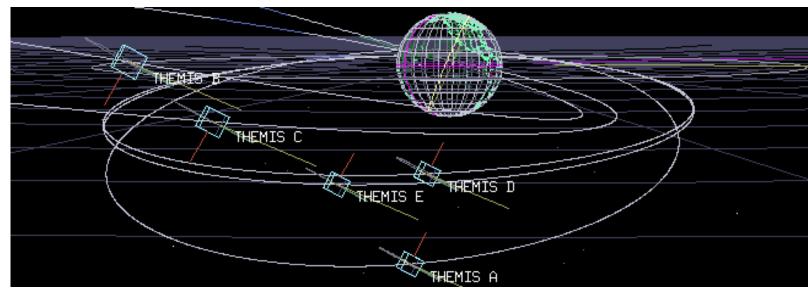






Flight Segment

- Constellation of 5 identical spacecraft
- Spin-stabilized at 20 rpm
- Identical suite of 5 science instruments to measure fields, waves and particles
- Hydrazine blow-down propulsion system with 49 kg of fuel and total ΔV capacity of 930 m/s, assuming 4 kg of fuel for attitude control
- S-band communications with 10 TLM data rates
- Two-way Doppler tracking and ranging for OD

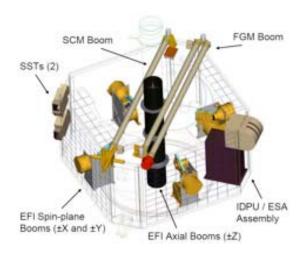


Orbit view on Dec. 6, 2008





Instrument Suite

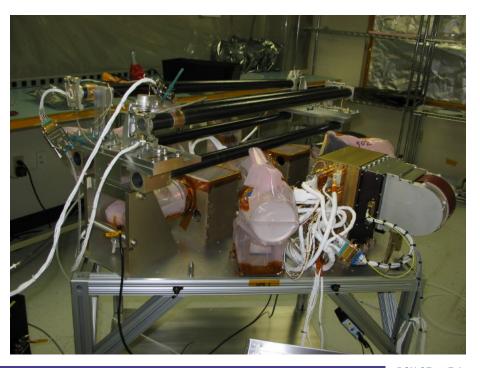


Particle Instruments (2)

- Electrostatic Analyzer (ESA) Low-energy ions and electrons
- Solid State Telescope (SST) High-energy ions and electrons

Fields and Waves Instruments (3)

- Fluxgate Magnetometer (FGM) DC and low-frequency magnetic fields
- Search Coil Magnetometer (SCM) AC magnetic fields
- Electric Field Instrument (EFI) Flectric fields and waves







Prime Mission Status

- Completed LEOP, coast phase, tail 1 and dayside 1 seasons
- Upcoming tail 2 and dayside 2 seasons
- Answered primary scientific questions already: What is the origin and time sequence of substorms?

Constellation Status

- All probes very healthy and collecting science data
- Fuel reserves allow for ambitious mission extension

Ground Systems & Operations Status

- All ground systems processing nominally
- Executed 253 individual thrust operations to date
- Completed 11,000 passes
- All routine passes supported in lights-out mode



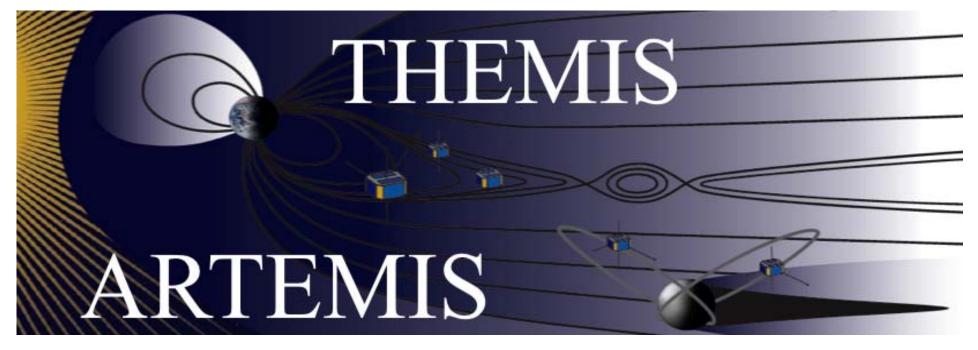


Time History of Events and Macroscale Interactions during

Substorms (THEMIS)

3 Probes (A/P5, D/P3, E/P4) in synchronized Earth orbits





Acceleration, Reconnection, Turbulence and Electrodynamics of the Moon's Interaction with the Sun (ARTEMIS)

2 Probes (B/P1, C/P2) in lunar orbits



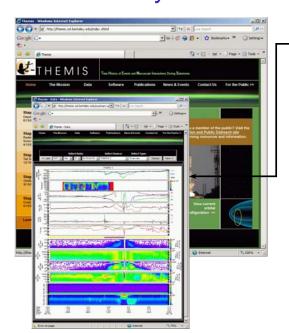


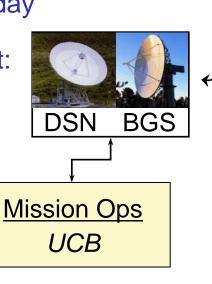
ARTEMIS Operations Concept

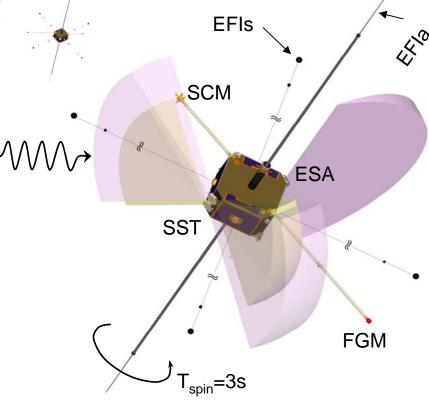
Lunar orbit periods: ~1 day Data volume: 100 Mbits/day

> **DSN** support: 3.5 hrs daily

Data products: 2 hr latency







Science instruments:

ESA: Electrostatic Analyzer (3 eV - 25 keV; *Carlson & McFadden*)

SST: Solid State Telescopes (25 - 1000 keV; *Larson*)

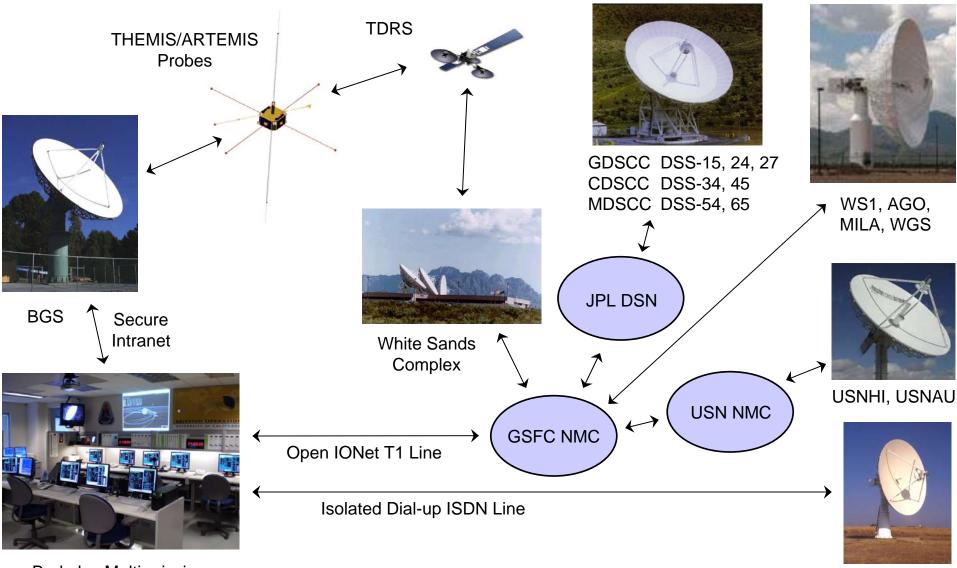
FGM: Fluxgate Magnetometer (0 - 128 Hz; Glassmeier, Auster & Baumjohann

SCM: Search Coil Magnetometer (1 Hz - 4 kHz; Roux & LeContel)

EFI: Electric Field Instrument (0 - 8 kHz; *Bonnell & Mozer*)







Berkeley Multi-mission **Operations Center**

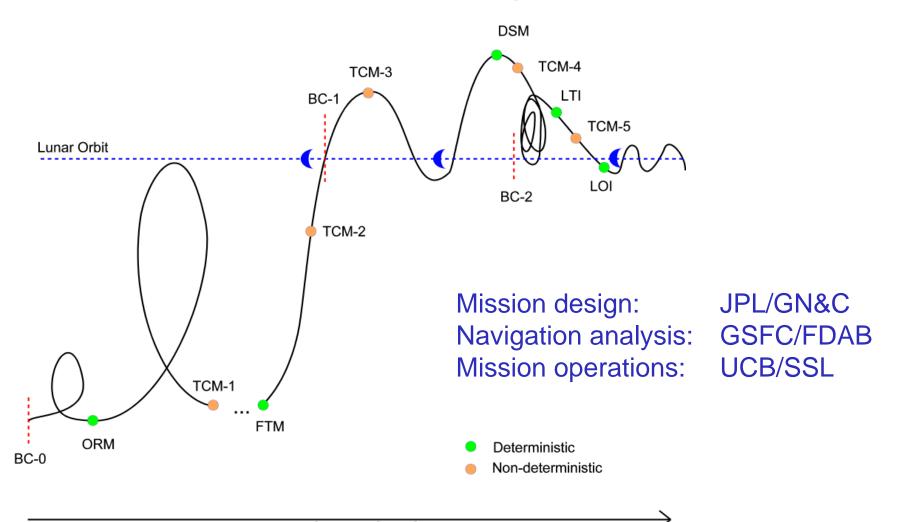
THEMIS/ARTEMIS Network Links

HBK





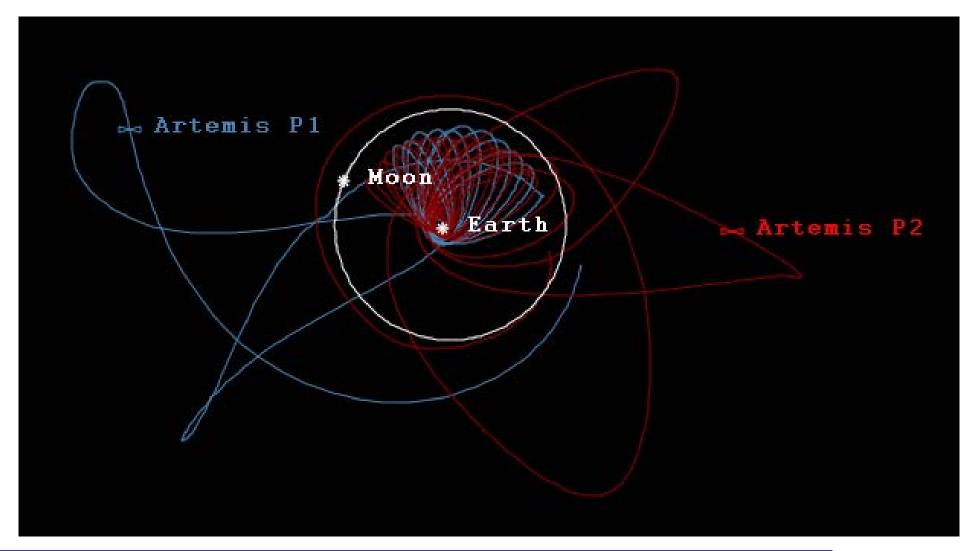
Lunar Transfer Scenario







ARTEMIS Trajectories

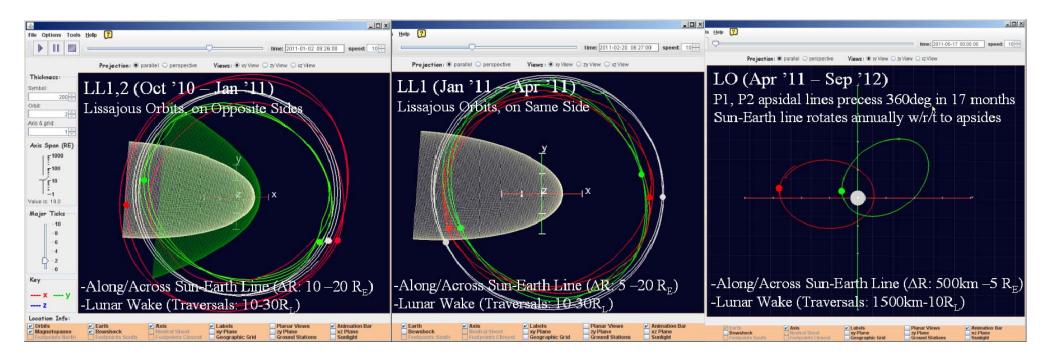






ARTEMIS Mission Phases

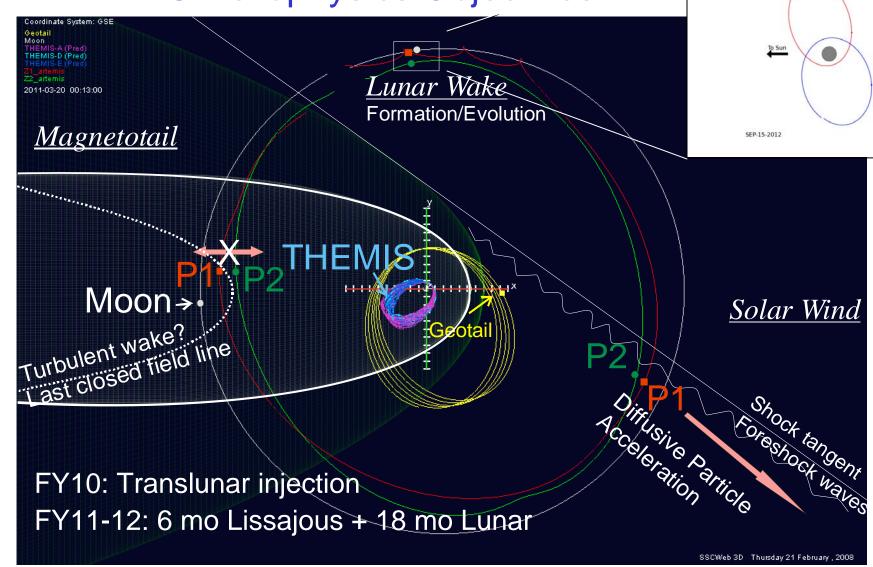
Insertion: FY10 Science: FY11 - FY12







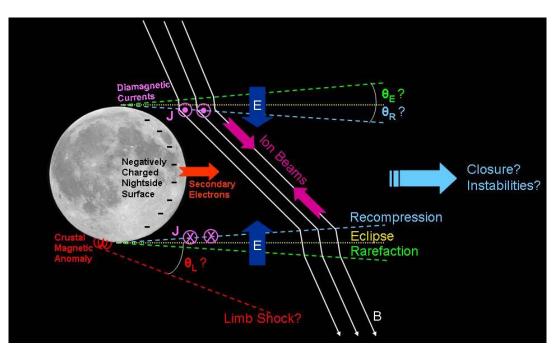
ARTEMIS Heliophysics Objectives

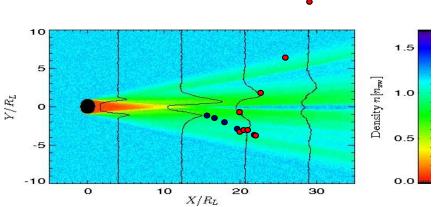






Electrodynamic Environment of the Moon: Lunar Wake



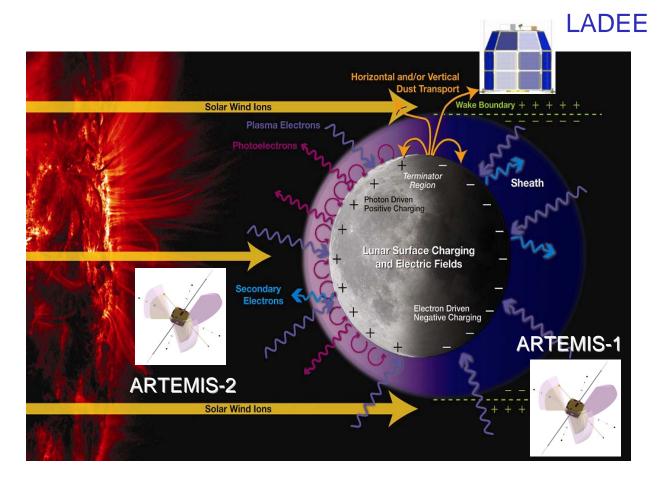


- The Moon is an ideal plasma laboratory for studying magnetized plasmas.
- The simple geometry of the Moon and its absorbing boundary produces a system that can be readily modeled with computers.
- This allows the software to be tested and the physics to be understood.
- ARTEMIS measurements are particularly useful for such tests.





Dust Levitation in Electric Fields



- APOLLO observers saw dust elevated above the lunar surface to possibly high altitudes.
- LADEE will probably have a dust detector, but has no capability of measuring the solar wind electric and magnetic fields.
- ARTEMIS measures the solar wind velocity, the interplanetary magnetic field and hence the solar wind electric field. It also can measure surface potentials with electron reflectometry.
- ARTEMIS plus LADEE will enable us to determine the response of charged lunar dust to the lunar and solar wind electric fields.

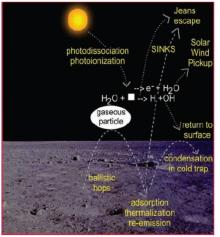




- Support of LRO
 - ARTEMIS can provide comprehensive monitoring of Lunar Space Environment
 - Complements LRO/CRATER providing measurements below 6 MeV
 - ARTEMIS has already supported LRO to certify ground stations (WS1, USN)

CRaTER Cosmic Ray Telescope for the Effects of Radation

- Support of NASA's Scientific Content of Exploration of the Moon
 - Understand the lunar atmosphere
- Support of all missions
 - ARTEMIS provides comprehensive monitoring of plasma conditions and lunar surface electric fields
 - Allows study of the response of the lunar exosphere and dust to external drivers
 - ARTEMIS provides solar wind monitoring



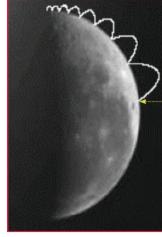
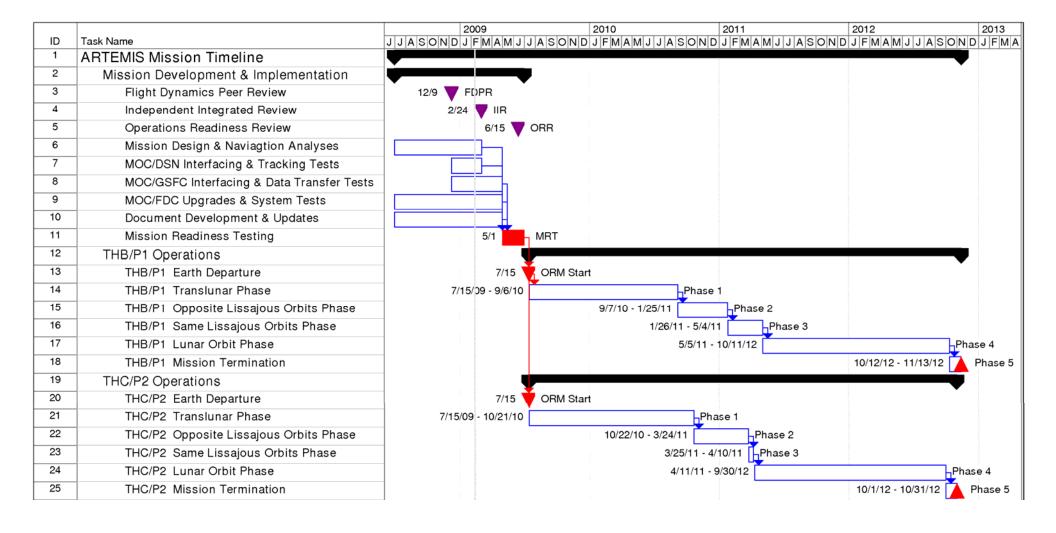


FIGURE 3.7 Lunar volatile transport. SOURCE: D.H. Crider and R.R. Vondrak, The solar wind as a possible source of lunar polar hydrogen deposits, J. Geophys. Res. 105:26773, 2000, ©2000 American Geophysical Union. Reproduced by permission of American Geophysical Union.





ARTEMIS Mission Timeline







Summary

- THEMIS prime mission accomplished by early **April 2009**
- ARTEMIS will provide exciting new science
- Lunar mission extension phase begins with orbit raise maneuvers in July 2009
- DSN interfacing and testing is on track
- Operations team is looking forward to continue work with DSN and appreciates the excellent support we already received